

①

HW 5.

1. Griffiths 4.45 p192

From wikipedia, the radius of the proton is  $R_p = 0.8775 \times 10^{-15} \text{ m}$   
 and the radius of the deuteron is  $R_d = 2.1424 \times 10^{-15} \text{ m}$ .

The Bohr radius  $a = 0.529 \times 10^{-10} \text{ m}$ .

Therefore the size of the hydrogen nucleus is about 4 or 5 orders of magnitude smaller than the expansion of the ground state of hydrogen, which has a size of the Bohr radius.

(a) The ground state wavefunction of H:

$$\Psi_{100}(r, \theta, \phi) = \frac{1}{\sqrt{\pi a^3}} e^{-r/a}$$

The probability of finding the electron inside a sphere of radius  $b$  is

$$\begin{aligned} P_{r < b} &= \int_0^b dr \int_0^\pi d\theta \int_0^{2\pi} d\phi |\Psi_{100}(r, \theta, \phi)|^2 r^2 \sin\theta \\ &= 2\pi \int_0^b dr \int_0^\pi d\theta \frac{1}{\pi a^3} e^{-2r/a} r^2 \sin\theta \\ &= \frac{2}{a^3} \left( \int_0^b dr r^2 e^{-2r/a} \right) \left( \int_0^\pi \sin\theta d\theta \right) \\ &= \frac{2}{a^3} \cdot \left( -\frac{1}{4} \right) a \cdot \left[ e^{-\frac{2r}{a}} (a^2 + 2ar + 2r^2) \right]_0^b \cdot 2 \\ &= -\frac{1}{a^2} \left[ e^{-\frac{2b}{a}} (a^2 + 2ab + 2b^2) - a^2 \right] \\ &= 1 - e^{-\frac{2b}{a}} \left( 1 + \frac{2b}{a} + \frac{2b^2}{a^2} \right) \end{aligned}$$

This is the exact answer.

(2)

$$(b) \epsilon \equiv \frac{2b}{a} \ll 1$$

$$\begin{aligned} P_{r < b} &= 1 - e^{-\frac{2b}{a}} \left( 1 + \frac{2b}{a} + \frac{2b^2}{a^2} \right) \\ &= 1 - e^{-\epsilon} \left( 1 + \epsilon + \frac{\epsilon^2}{2} \right) \\ &\approx 1 - (1-\epsilon)(1+\epsilon) \\ &\approx 1 - \left( 1 - \epsilon + \frac{\epsilon^2}{2} - \frac{\epsilon^3}{6} \right) \left( 1 + \epsilon + \frac{\epsilon^2}{2} \right) \\ &= 1 - \left[ 1 - \cancel{\epsilon} + \frac{\epsilon^2}{2} - \frac{\epsilon^3}{6} + \cancel{\epsilon} - \epsilon^2 + \frac{\epsilon^3}{2} - \frac{\epsilon^4}{6} + \frac{\epsilon^2}{2} - \cancel{\frac{\epsilon^3}{2}} + \frac{\epsilon^4}{4} - \frac{\epsilon^5}{12} \right] \\ &= 1 - \left[ 1 - \frac{\epsilon^3}{6} + \frac{\epsilon^4}{12} + \dots \right] \\ &= \frac{1}{6} \epsilon^3 + O(\epsilon^4) \\ &\approx \frac{1}{6} \epsilon^3 = \frac{1}{6} \left( \frac{2b}{a} \right)^3 = \frac{4}{3} \left( \frac{b}{a} \right)^3 \end{aligned}$$

The lowest order is  $\epsilon^3$ , or  $(\frac{b}{a})^3$ ,  $P_{r < b} \approx \frac{4}{3} \left( \frac{b}{a} \right)^3$

(c) If we assume that  $\psi(r)$  is essentially constant over the tiny sphere with radius  $b$ , then  $P_{r < b}$  can be obtained as

$$P_{r < b} \approx \frac{4}{3} \pi b^3 \cdot |\psi(r=0)|^2 = \frac{4}{3} \pi b^3 \cdot \frac{1}{\pi a^3} = \frac{4}{3} \left( \frac{b}{a} \right)^3$$

which is the same as the result in (b).

$$(d) b \approx 10^{-15} \text{ m}, a \approx 0.5 \times 10^{-10} \text{ m}$$

$$P_{r < b} \approx \frac{4}{3} \left( \frac{b}{a} \right)^3 = \frac{4}{3} \cdot \left( \frac{10^{-15}}{0.5 \times 10^{-10}} \right)^3 = \frac{32}{3} \times 10^{-15} = 1.07 \times 10^{-14}$$

which is an extremely small probability.

2. First we want to prove that  $L_+ Y_e^m = \hbar \sqrt{\ell(\ell+1) - m(m+1)} Y_e^{m+1}$

This is also problem 4.18 of Griffiths.

Let us assume  $L_+ Y_e^m = A Y_e^{m+1}$  and  $L_- Y_e^m = B Y_e^{m-1}$

From [4.112] of Griffiths

$$L^2 = L_+ L_- + L_z^2 - \hbar L_z = L_- L_+ + L_z^2 + \hbar L_z \quad (*)$$

let us first use the latter relation, from which

$$\begin{aligned} L_- L_+ &= L^2 - L_z^2 - \hbar L_z \\ \Rightarrow L_- L_+ Y_e^m &= L^2 Y_e^m - L_z^2 Y_e^m - \hbar L_z Y_e^m = [\hbar^2 \ell(\ell+1) - \hbar^2 m^2 - \hbar^2 m] Y_e^m \\ &= \hbar^2 [\ell(\ell+1) - m(m+1)] Y_e^m \end{aligned}$$

Now multiply  $(Y_e^m)^*$  on the left for both sides and then take an integration over angle

$$\Rightarrow \int (Y_e^m)^* L_- L_+ Y_e^m d\Omega = \hbar^2 [\ell(\ell+1) - m(m+1)] \int (Y_e^m)^* Y_e^m d\Omega$$

use  $(AB)^+ = B^+ A^+$  and  $L_+^+ = L_-$ ,  $L_-^+ = L_+$

$$\Rightarrow \int (L_+ Y_e^m)^* L_+ Y_e^m d\Omega = \hbar^2 [\ell(\ell+1) - m(m+1)]$$

$$\Rightarrow \int \cancel{(Y_e^m)^*} A^* A Y_e^m d\Omega = \hbar^2 [\ell(\ell+1) - m(m+1)]$$

$$\Rightarrow |A|^2 \underbrace{\int (Y_e^m)^* Y_e^m d\Omega}_{=1} = \hbar^2 [\ell(\ell+1) - m(m+1)]$$

$$\Rightarrow A = \hbar \sqrt{\ell(\ell+1) - m(m+1)}$$

Similarly, if you use the first relation of (\*) you will get

$$B = \hbar \sqrt{\ell(\ell+1) - m(m-1)}$$

$$\begin{cases} L_+ = L_x + i L_y \\ L_- = L_x - i L_y \end{cases} \Rightarrow \begin{cases} L_x = \frac{1}{2} (L_+ + L_-) \\ L_y = \frac{1}{2i} (L_+ - L_-) \end{cases}$$

$$\begin{aligned}
 & \langle Y_{lm_1} | L_x | Y_{lm_2} \rangle \\
 &= \frac{1}{2} \langle Y_{lm_1} | L_+ | Y_{lm_2} \rangle + \frac{1}{2} \langle Y_{lm_1} | L_- | Y_{lm_2} \rangle \\
 &= \frac{1}{2} \langle Y_{lm_1} | \hbar \sqrt{\ell(\ell+1)-m_2(m_2+1)} | Y_{l,m_2+1} \rangle \\
 &\quad + \frac{1}{2} \langle Y_{lm_1} | \hbar \sqrt{\ell(\ell+1)-m_2(m_2-1)} | Y_{l,m_2-1} \rangle \\
 &= \frac{\hbar}{2} \left[ \delta_{m_1, m_2+1} \sqrt{\ell(\ell+1)-m_2(m_2+1)} + \delta_{m_1, m_2-1} \sqrt{\ell(\ell+1)-m_2(m_2-1)} \right]
 \end{aligned}$$

$$\begin{aligned}
 & \langle Y_{lm_1} | L_y | Y_{lm_2} \rangle \\
 &= \frac{1}{2i} \left[ \langle Y_{lm_1} | L_+ | Y_{lm_2} \rangle - \langle Y_{lm_1} | L_- | Y_{lm_2} \rangle \right] \\
 &= \frac{1}{2i} \left[ \hbar \sqrt{\ell(\ell+1)-m_2(m_2+1)} \langle Y_{lm_1} | Y_{l,m_2+1} \rangle \right. \\
 &\quad \left. - \hbar \sqrt{\ell(\ell+1)-m_2(m_2-1)} \langle Y_{lm_1} | Y_{l,m_2-1} \rangle \right] \\
 &= \frac{\hbar}{2i} \left[ \delta_{m_1, m_2+1} \sqrt{\ell(\ell+1)-m_2(m_2+1)} - \delta_{m_1, m_2-1} \sqrt{\ell(\ell+1)-m_2(m_2-1)} \right]
 \end{aligned}$$

(4)

$$3. H = \frac{L_x^2 + L_y^2}{2I_1} + \frac{L_z^2}{2I_2} = \frac{L^2 - L_z^2}{2I_1} + \frac{L_z^2}{2I_2} = \frac{1}{2I_1} L^2 + \frac{I_1 - I_2}{2I_1 I_2} L_z^2$$

Because  $Y_l^m(\theta, \phi)$  are eigenstates of  $L^2$  and  $L_z$ , they are also the eigenstates of  $H$ .

The eigenvalues of  $H$  are

$$\frac{1}{2I_1} \hbar^2 l(l+1) + \frac{I_1 - I_2}{2I_1 I_2} \hbar^2 m^2 = \alpha l(l+1) + \beta m^2$$

where we have set  $\alpha = \frac{\hbar^2}{2I_1}$  and  $\beta = \frac{I_1 - I_2}{2I_1 I_2} \hbar^2 > 0$  (b/c  $I_1 > I_2$ )

$$l = 0, 1, 2, \dots \quad m = -l, -l+1, \dots, l-1, l$$

Suppose  $\beta < \alpha$ ,  $\Rightarrow \frac{I_1 - I_2}{2I_1 I_2} \hbar^2 < \frac{\hbar^2}{2I_1}$ ,  $I_1 < 2I_2$  (i.e.,  $I_2 < I_1 < 2I_2$ )

then the first few eigenlevels are sketched below

$$12\alpha + 9\beta \text{ --- } (l, m) = (3, 3) \text{ or } (3, -3)$$

$$12\alpha + 4\beta \text{ --- } (l, m) = (3, 2) \text{ or } (3, -2)$$

$$12\alpha + \beta \text{ --- } (l, m) = (3, 1) \text{ or } (3, -1)$$

$$12\alpha \text{ --- } (l, m) = (3, 0)$$

$$6\alpha + 4\beta \text{ --- } (l, m) = (2, 2) \text{ or } (2, -2)$$

$$6\alpha + \beta \text{ --- } (l, m) = (2, 1) \text{ or } (2, -1)$$

$$6\alpha \text{ --- } (l, m) = (2, 0)$$

$$2\alpha + \beta \text{ --- } (l, m) = (1, 1) \text{ or } (1, -1)$$

$$2\alpha \text{ --- } (l, m) = (1, 0)$$

$$0 \text{ --- } (l, m) = (0, 0)$$

If  $\beta$  is very big, then the ordering of the levels can be complicated and this case will not be considered here.

(5)

4. In spherical coordinates

$$\begin{cases} x = r \sin\theta \cos\phi \\ y = r \sin\theta \sin\phi \\ z = r \cos\theta \end{cases}$$

$$\begin{aligned}\psi(x, y, z) &= C(xy + yz + zx) e^{-dr^2} \\ &= C r^2 e^{-dr^2} (\sin^2\theta \sin\phi \cos\phi + \sin\theta \cos\theta \sin\phi + \sin\theta \cos\theta \cos\phi)\end{aligned}$$

Because angular momentum is determined only by the angular part of the wavefunction, we shall only focus on the angular part:  $Y(\theta, \phi) = A(\sin^2\theta \sin\phi \cos\phi + \sin\theta \cos\theta \sin\phi + \sin\theta \cos\theta \cos\phi)$

where  $A$  is the normalization constant.

The normalization condition is

$$\begin{aligned}1 &= \int_0^{2\pi} \int_0^\pi |Y(\theta, \phi)|^2 \sin\theta d\theta d\phi \\ &= |A|^2 \int_0^{2\pi} \int_0^\pi (\sin^2\theta \sin\phi \cos\phi + \sin\theta \cos\theta \sin\phi + \sin\theta \cos\theta \cos\phi)^2 \sin\theta d\theta d\phi \\ &= |A|^2 \cdot \frac{4\pi}{5} \quad \Rightarrow \quad A = \sqrt{\frac{5}{4\pi}}\end{aligned}$$

The eigenstates of the angular wavefunction are spherical harmonics. Therefore  $Y(\theta, \phi)$  can be written as a linear combination of spherical harmonics

$$Y(\theta, \phi) = \sum_{lm} C_l^m Y_l^m(\theta, \phi)$$

$$\text{where } C_l^m = \langle Y_l^m | Y \rangle = \int_0^{2\pi} \int_0^\pi [Y_l^m]^* Y \sin\theta d\theta d\phi$$

(6)

$Y_l^m$  are eigenstates of  $L^2$  and  $L_z$  with

$$L^2 Y_l^m = \hbar^2 l(l+1) Y_l^m \quad L_z Y_l^m = \hbar m Y_l^m$$

The probability that a measurement on  $L^2$  yields 0 (i.e.,  $l=0, m=0$ )

is  $|C_0|^2 = |\langle Y_0 | Y \rangle|^2$

$$Y_0 = \sqrt{\frac{1}{4\pi}}$$

$$= \left| \int_0^{2\pi} d\phi \int_0^\pi d\theta \sqrt{\frac{1}{4\pi}} \cdot \sqrt{\frac{5}{4\pi}} (\sin^2 \theta \sin \phi \cos \phi + \sin \theta \cos \theta \sin \phi + \sin \theta \cos \theta \cos \phi) \sin \theta \right|^2$$

$$= 0$$

The probability that a measurement on  $L^2$  yields  $6\hbar^2$  (i.e.,  $l=2$ ) is

$$|C_2^m|^2 = |C_2^{-2}|^2 + |C_2^{-1}|^2 + |C_2^0|^2 + |C_2^1|^2 + |C_2^2|^2$$

Using Mathematica, I get

$$C_2^{-2} = -\frac{i}{\sqrt{6}} \Rightarrow |C_2^{-2}|^2 = \frac{1}{6}$$

$$C_2^{-1} = \frac{1-i}{\sqrt{6}} \Rightarrow |C_2^{-1}|^2 = \frac{1}{3}$$

$$C_2^0 = 0 \Rightarrow |C_2^0|^2 = 0$$

$$C_2^1 = -\frac{1+i}{\sqrt{6}} \Rightarrow |C_2^1|^2 = \frac{1}{3}$$

$$C_2^2 = \frac{i}{\sqrt{6}} \Rightarrow |C_2^2|^2 = \frac{1}{6}$$

$$\Rightarrow |C_2^m|^2 = \frac{1}{6} + \frac{1}{3} + 0 + \frac{1}{3} + \frac{1}{6} = 1$$

(7)

$$5. (a) |jm\rangle = \sum_{m_1=-j_1}^{j_1} \sum_{m_2=-j_2}^{j_2} \underbrace{\langle j_1 j_2; m_1 m_2 | jm \rangle}_{\text{C-G coefficients}} |j_1 m_1\rangle |j_2 m_2\rangle$$

There are a lot of places to find the C-G coefficients. If you know how to read the C-G coefficient table (Table 4.8) on page 188, you can find them there. Besides the traditional C-G tables, I find that there are some on-line C-G calculators which are extremely straightforward to use, for example, the WolframAlpha Clebsch-Gordan calculator. You just input the values of  $j_1, j_2, m_1, m_2, j, m$  in the corresponding boxes and you get the C-G coefficient out.

You can check that  $\frac{1}{\sqrt{5}}, \sqrt{\frac{3}{5}}, \frac{1}{\sqrt{5}}$  are indeed the correct C-G coefficients.

$$|30\rangle = \frac{1}{\sqrt{5}} |21\rangle |1-1\rangle + \sqrt{\frac{3}{5}} |20\rangle |10\rangle + \frac{1}{\sqrt{5}} |2-1\rangle |11\rangle$$

$$\begin{array}{ccccccccc} \uparrow & \uparrow \\ j_1 & m_1 & j_2 & m_2 & j_1 & m_1 & j_2 & m_2 & j_1 & m_1 & j_2 & m_2 \end{array}$$

(b) Strictly speaking,  $|20\rangle$  or  $|10\rangle$  can be expanded in many different ways depending on the value of  $j_1$  and  $j_2$ . Here we follow the way that  $|30\rangle$  was expanded, i.e.,  $j_1=2, j_2=1$ . Then

$$|20\rangle = \frac{1}{\sqrt{2}} |21\rangle |1-1\rangle + 0 \cdot |20\rangle |10\rangle - \frac{1}{\sqrt{2}} |2-1\rangle |11\rangle$$

$$|10\rangle = \sqrt{\frac{3}{10}} |21\rangle |1-1\rangle - \sqrt{\frac{2}{5}} |20\rangle |10\rangle + \sqrt{\frac{3}{10}} |2-1\rangle |11\rangle$$

(8)

$$\begin{aligned}
 (c) \langle 10|20 \rangle &= \left( \sqrt{\frac{3}{10}} \langle 21| \langle 1-1| - \sqrt{\frac{2}{5}} \langle 20| \langle 10| + \sqrt{\frac{3}{10}} \langle 2-1| \langle 11| \right) \\
 &\quad \cdot \left( \frac{1}{\sqrt{2}} |21\rangle |1-1\rangle + 0 \cdot |20\rangle |10\rangle - \frac{1}{\sqrt{2}} |2-1\rangle |11\rangle \right) \\
 &= \left[ \sqrt{\frac{3}{10}} \cdot \frac{1}{\sqrt{2}} - \sqrt{\frac{2}{5}} \cdot \frac{1}{\sqrt{2}} \right] = 0
 \end{aligned}$$

$$\begin{aligned}
 \langle 10|30 \rangle &= \left( \sqrt{\frac{3}{10}} \langle 21| \langle 1-1| - \sqrt{\frac{2}{5}} \langle 20| \langle 10| + \sqrt{\frac{3}{10}} \langle 2-1| \langle 11| \right) \\
 &\quad \cdot \left( \frac{1}{\sqrt{5}} |21\rangle |1-1\rangle + \sqrt{\frac{3}{5}} |20\rangle |10\rangle + \frac{1}{\sqrt{5}} |2-1\rangle |11\rangle \right) \\
 &= \left[ \sqrt{\frac{3}{10}} \cdot \frac{1}{\sqrt{5}} - \sqrt{\frac{2}{5}} \cdot \sqrt{\frac{3}{5}} + \sqrt{\frac{3}{10}} \cdot \frac{1}{\sqrt{5}} \right] = 0
 \end{aligned}$$

$$\begin{aligned}
 \langle 20|30 \rangle &= \left( \frac{1}{\sqrt{2}} \langle 21| \langle 1-1| + 0 \cdot \langle 20| \langle 10| - \frac{1}{\sqrt{2}} \langle 2-1| \langle 11| \right) \\
 &\quad \cdot \left( \frac{1}{\sqrt{5}} |21\rangle |1-1\rangle + \sqrt{\frac{3}{5}} |20\rangle |10\rangle + \frac{1}{\sqrt{5}} |2-1\rangle |11\rangle \right) \\
 &= \frac{1}{\sqrt{2}} \cdot \frac{1}{\sqrt{5}} - \frac{1}{\sqrt{2}} \cdot \frac{1}{\sqrt{5}} = 0
 \end{aligned}$$

So  $|10\rangle$ ,  $|20\rangle$ ,  $|30\rangle$  are orthogonal to each other.

## 6. 3D isotropic harmonic oscillator

(a)  $V(r) = \frac{1}{2}m\omega^2 r^2$

Because the potential is spherically symmetric, the Schrödinger equation can be handled by separation of variables in spherical coordinates. The wavefunction is  $\psi_{nlm}(r, \theta, \phi) = R_n(r) Y_l^m(\theta, \phi)$ .

where the radial part of the wavefunction  $R_{nl}(r)$  satisfies the following equation

$$\left[ -\frac{\hbar^2}{2m} \frac{1}{r} \frac{d^2}{dr^2} r + \frac{1}{2}m\omega^2 r^2 + \frac{l(l+1)\hbar^2}{2mr^2} \right] R_{nl}(r) = E_{nl} R_{nl}(r)$$

It is quite tricky to work this radial equation out. Here I just write down the solution

$$R_{nrl} = \sum_{k=0}^{\infty} a_k y^{l+2k} e^{-y^2/2}$$

$$\text{where } y = \sqrt{\frac{m\omega}{\hbar}} r$$

$$\text{and } E_{nrl} = \left( 2n_r + l + \frac{3}{2} \right) \hbar\omega$$

$$n_r = 0, 1, 2, \dots \quad l = 0, 1, 2, \dots$$

(b) The Schrödinger equation of the 3D isotropic harmonic oscillator in Cartesian coordinates is

$$-\frac{\hbar^2}{2m} \left( \frac{\partial^2 \Psi}{\partial x^2} + \frac{\partial^2 \Psi}{\partial y^2} + \frac{\partial^2 \Psi}{\partial z^2} \right) + \frac{1}{2}m\omega^2(x^2 + y^2 + z^2) \Psi = E\Psi$$

Let  $\Psi(x, y, z) = X(x)Y(y)Z(z)$ , plug into the above equation

$$-\frac{\hbar^2}{2m} \left( \frac{d^2 X}{dx^2}YZ + \frac{d^2 Y}{dy^2}ZX + \frac{d^2 Z}{dz^2}XY \right)$$

$$+ \frac{1}{2}m\omega^2(x^2 + y^2 + z^2)XYZ = EXYZ$$

divide XYZ both sides

$$\left( -\frac{\hbar^2}{2m} \frac{1}{X} \frac{d^2X}{dx^2} + \frac{1}{2} m\omega^2 x^2 \right) + \left( -\frac{\hbar^2}{2m} \frac{1}{Y} \frac{d^2Y}{dy^2} + \frac{1}{2} m\omega^2 y^2 \right) \\ + \left( -\frac{\hbar^2}{2m} \frac{1}{Z} \frac{d^2Z}{dz^2} + \frac{1}{2} m\omega^2 z^2 \right) = E$$

The first term is a function only of  $x$ , the second term only of  $y$ , and the third term only of  $z$ . For the above equation to hold, each term should equal a constant:

$$\left\{ \begin{array}{l} -\frac{\hbar^2}{2m} \frac{d^2X}{dx^2} + \frac{1}{2} m\omega^2 x^2 X = E_x X \quad \leftarrow \text{1D harmonic oscillator} \\ -\frac{\hbar^2}{2m} \frac{d^2Y}{dy^2} + \frac{1}{2} m\omega^2 y^2 Y = E_y Y \quad \leftarrow \text{1D } \dots \\ -\frac{\hbar^2}{2m} \frac{d^2Z}{dz^2} + \frac{1}{2} m\omega^2 z^2 Z = E_z Z \quad \leftarrow \text{1D } \dots \end{array} \right.$$

and  $E = E_x + E_y + E_z$

we know that

$$E_x = (n_x + \frac{1}{2}) \hbar\omega$$

$$n_x = 0, 1, 2, \dots$$

$$E_y = (n_y + \frac{1}{2}) \hbar\omega$$

$$n_y = 0, 1, 2, \dots$$

$$E_z = (n_z + \frac{1}{2}) \hbar\omega$$

$$n_z = 0, 1, 2, \dots$$

$$\Rightarrow E = E_x + E_y + E_z = (n_x + n_y + n_z + \frac{3}{2}) \hbar\omega$$

$$\equiv (n + \frac{3}{2}) \hbar\omega \quad \text{with } n = n_x + n_y + n_z$$

3D harmonic oscillator is decomposed into 3 1D harmonic oscillators

(c) The first 4 energy levels are

From spherical coordinates

E	$n_x n_y n_z$	degeneracy	$n_r l$	degeneracy = $(2l+1)$
$\frac{3}{2}\hbar\omega$	0 0 0	1	0 0	1
$\frac{5}{2}\hbar\omega$	1 0 0 0 1 0 0 0 1	3	0 1	3
$\frac{7}{2}\hbar\omega$	2 0 0 0 2 0 0 0 2 1 1 0 1 0 1 0 1 1	6	1 0 0 2	$1+5=6$
$\frac{9}{2}\hbar\omega$	3 0 0 0 3 0 0 0 3 2 1 0 2 0 1 1 2 0 0 2 1 1 0 2 0 1 2 1 1 1	10	1 1 0 3	$3+7=10$