Simplified version of the random-phase-approximation-with-exchange method*

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By comparing the random-phase-approximation-with-exchange (RPAE) method of Amusia et al. for the calculation of photoionization cross sections from closed-shell atoms with the reaction-matrix method, we obtain a simplified version of the RPAE by which we calculate the dipole-transition amplitude much as in the reaction-matrix method, yet with ground-state correlation effects included simultaneously. Calculation of photoionization cross sections for the \( 3p \rightarrow \epsilon d \) transition of Ar and \( 4d \rightarrow \epsilon f \) transition of Xe by this method yields a substantial improvement on the calculation by the reaction-matrix method and good agreement with RPAE and with experiment.

I. INTRODUCTION

The photoionization spectra of the noble gases in the far ultraviolet have been studied in the last ten years experimentally and theoretically and have served as critical tests of theoretical procedures. For He and Ne, single-electron photoionization is reproduced fairly well by the independent-particle approximation.\(^1\) For Ar, Kr, and Xe, on the other hand, qualitative discrepancies appeared between experiment and the result of independent particle approximations. Moreover, these discrepancies were only reduced but not eliminated by calculations that improved only the continuum final-state wave functions.\(^2\) A good qualitative and quantitative fit of these spectra was first achieved by Amusia et al.\(^3\) through RPAE calculations (random-phase approximation with exchange) which introduce a simultaneous improvement on the wave functions of the ground state and the continuum states. Similar results have been obtained by Wendin.\(^4\)

This paper presents a simplified version of the RPAE method (SRPAE), which allows one to calculate the photoionization cross section in a way similar to the more familiar configuration-interaction procedures, specifically to the reaction-matrix method of Ref. 1, yet including an improvement of the ground-state wave function. In Sec. II, after a brief review of the reaction-matrix method and the RPAE method, we discuss the similarities and differences of the two approaches. It is then shown how to develop a SRPAE method by making approximations on the RPAE equations. The method utilizes an approximate property of interaction-matrix elements, which has been previously applied by Wendin.\(^1\) The formulas thus obtained have a structure similar to those of the reaction-matrix method, but with extra terms resulting from improving the ground-state wave function simultaneously. In Sec. III, we use SRPAE method to re-calculate the photoionization cross section for the \( 3p \rightarrow \epsilon d \) transition of Ar and the \( 4d \rightarrow \epsilon f \) transition of Xe. In fact, this calculation is done using the same residual interaction-matrix elements as were used by Starace\(^5\) in his calculation by the reaction-matrix method. The substantial improvement of the cross sections calculated by this method, demonstrated by the curves in Figs. 1 and 2, manifests the importance of ground-state correlation effect. The result is also compared with the RPAE calculations of Amusia et al.\(^3\) and with experiment.

II. SIMPLIFIED VERSION

OF RPAE METHOD

The reaction-matrix method\(^5,6\) represents the final-state wave function of the photoelectron-plasmon system as

\[
\psi_E = \sum_i \left( u_{IE} B_i(E) \right) + \sum_j P \int d\epsilon \frac{u_{iE} K_{ie,JE}(E)}{E - \epsilon} B_j(E). \tag{1}
\]

In this equation, \( u_{iE} \) is a Slater-determinant basis function of the many-electron system in an independent electron approximation, with \( \epsilon \) the total energy of the system, \( i \) the channel index represented by an appropriate set of discrete quantum numbers; \( P \) indicates principal-part integration over the continuum states and summation over the discrete states. The coefficients \( B \) in Eq. (1) are determined by boundary conditions.\(^5,6\) The matrix elements \( K_{ie,JE}(E) \) are obtained by solving

\[
K_{ie,JE}(E) = V_{ie,JE} + \sum_l \int d\epsilon' V_{iE,ie'} \times \left[ P/(E - \epsilon') \right] K_{ie',JE}(E), \tag{2}
\]
with
\[ V_{\text{res}} = (u_{i\xi}, V_{\text{res}} u_{j\eta}), \]  
\[ (3) \]

where
\[ V_{\text{res}} = -\sum_i \left( \frac{Z_i}{r_i} + V_{\text{res}}(r_i) \right) + \sum_{i<j} \frac{1}{r_{ij}}, \]  
\[ (4) \]
is the residual interaction that was disregarded in generating the basis-set wave functions with model potential \( V_{\text{res}}(r) \); the summation in (4) extends over all the electrons in the atom with nuclear charge \( Z \).

If we assume that the core electrons remain unresolved during the photoionization process, the single-particle wave functions of the core electrons remain the same for the initial- and final-state wave functions of the system. Under this approximation, Eq. (1) constitutes essentially an improvement of the photoelectron wave function. If we use an "unimproved" ground-state wave function, i.e., a Slater determinant \( \psi_0 \) of single-particle wave functions in the ground-state configuration, then the dipole amplitude \( D^{(1)}(E) \) for transition to the channel \( i \) is given by

\[ D^{(1)}(E) = \sum_i \left( d_{i\xi} B_i(E) + \sum_j P \int d\epsilon \frac{d\epsilon}{\epsilon} \frac{K_{i\xi j\eta}(E)}{\epsilon - \epsilon} B_j(E) \right), \]  
\[ (5) \]

where \( d_{i\xi} = (\psi_0 | \epsilon_i | u_{i\xi}) \), with \( j \) summed over all the electrons.

In this method, one usually generates the basis function by solving a Schrödinger equation with the model potential \( V_{\text{res}}(r) \) tabulated by Herman and Skillman. These basis functions are then used to evaluate the residual interaction-matrix elements \( V_{i\xi j\eta} \) in Eq. (3) and the dipole amplitudes \( d_{i\xi} \). After Eq. (2) is solved one can find the dipole amplitude \( D^{(1)}(E) \) by doing the principal-part integration in Eq. (5).

The formulation of the RPA can be found in a number of references and is not discussed here. In order to illustrate the close similarity between the equations of RPA and of the reaction-matrix method, we start from the formulas of Ref. 3 and use the name RPAE instead of RPA. Matrix elements analogous (but not identical) to the \( K \)-matrix elements of Eq. (2) are called \((n_i l_{i1}, n_j l_{j2}, \Gamma_i(\omega)) | n_j l_{j2}, n_i l_{i1} \rangle \) by Amusia et al. The matrix elements satisfy the system of equations

\[ (n_i l_{i1}, n_j l_{j2}, U_{i1} | n_j l_{j2}, n_i l_{i1}) = (n_i l_{i1}, n_j l_{j2}, U_{i1} | n_j l_{j2}, n_i l_{i1}), \] 
\[ (6) \]

which coincides with Eq. (18) of Ref. 3, except for correction of several misprints. In Eq. (6), \((n_i l_{i1}, n_j l_{j2}, U_{i1} | n_j l_{j2}, n_i l_{i1}) \) represents certain terms of the matrix element of the Hamiltonian between states in which two electrons have quantum numbers: \((n_i l_{i1}, n_j l_{j2}) \) and \((n_j l_{j2}, n_i l_{i1}) \), respectively; these terms include primarily the effects of 2p pole interaction between these electrons. Also, \( \delta \) is an infinitesimal real positive number, the summation is carried over all the intermediate single-particle states that are present \((n_a = F) \) or absent \((n_a > F) \) in the ground-state configuration of an atom with closed-shell structure. The summation is understood to extend over all hole states and all discrete particle states and to include integration over all continuum particle states. The Fermi step function \( n_a \) is defined as

\[ n_a = \begin{cases} 1, & n_a = F; \\ 0, & n_a > F, \end{cases} \]  
\[ (7) \]
and the matrix elements \( \langle n_1 l_1, n_2 l_2 \| U_{ij} \| n_3 l_3, n_4 l_4 \rangle \) are defined as in Eq. (19) of Ref. 3.

The radial part of the dipole transition amplitude in the RPAE formulation is given by

\[
\langle n_1 l_1 \| d \| n_2 l_2 \rangle + \left( \sum_{n_3 \neq n_4} \sum_{\frac{n_3 + \frac{l_1}{2}}{n_4 + \frac{l_1}{2}}} \right)^{\frac{1}{3}} \left( \begin{array}{ccc} 1 & 1 & 1 \\ 0 & 0 & 0 \\ 1 & 1 & 1 \end{array} \right) \left( \begin{array}{c} (2l_1 + 1) \omega - E_{n_1} + i\delta(1 - 2n_3) \\ (2l_2 + 1) \omega - E_{n_2} + i\delta \end{array} \right),
\]

which is identical to Eq. (21) of Ref. 3, except for correction of several misprints. The radial dipole integral is defined as

\[
\langle n_1 l_1 \| d \| n_2 l_2 \rangle = \int_0^\infty P_{n_1 l_1}(r) P_{n_2 l_2}(r) dr,
\]

with \( P_{n_1 l_1}(r) \) the radial wave function of the electron.

In the RPAE method, a Hartree-Fock basis was used to evaluate the dipole integral and the interaction-matrix elements \( \langle n_1 l_1, n_2 l_2 \| U_{ij} \| n_3 l_3, n_4 l_4 \rangle \). After Eq. (6) is solved for the \( \Gamma_i(\omega) \) matrix elements, one finds the radial dipole amplitude by doing the integration in Eq. (8). In the numerical calculations, a proper choice of the Hartree-Fock basis functions is required. This is explained by Amusia et al., in Ref. 3 and is not discussed here. It is to be noted here that Eq. (6) holds for arbitrary choice of the initial potential, irrespective of the fact that a Hartree-Fock basis affords a better first approximation. Thus Eq. (6) can still be used when one starts from a central field approximation, like the Hermann-Skillman's tabulated potential, with appropriate modification in the definition of the matrix elements of \( U_{ij} \). As we have said, the reaction-matrix method provides a way to calculate the dipole transition amplitude \( D^{(l)}(E) \) in Eq. (5), in which the final-state wave function is improved, whereas RPAE provides a method to calculate the dipole amplitude in Eq. (8) in which the final-state and the ground-state wave function are improved simultaneously. Roughly speaking, the difference in the equation of the two formulations should be due to the improvement of the ground-state wave function. Indeed Eqs. (6) and (8) are seen to include terms which have no analog in Eqs. (2) and (5).

In order to clarify the notations we specialize the general formulas (2), (5), (6), and (8) to the single-channel 3p–Ed transition of Ar. For this dipole transition, we set \( l = 1 \) in Eq. (6). The factor \( (2l + 1) \) in the denominator of Eq. (6) can be eliminated by defining \( \Gamma_i(\omega) = \Gamma_i(\omega)/(2l + 1) \) and \( U_{ij} = U_{ij}/(2l + 1) \). Specializing the hole state in Eq. (3) to 3p hole and the particle states to ed states, including nd with \( n > 3 \), Eq. (8) takes the form

\[
(\text{Ed} \| d \| 3p) = \int d\epsilon \frac{\epsilon d\epsilon}{\omega - \epsilon - E_{3p} + i\delta} \frac{\epsilon d\epsilon}{\omega - \epsilon - E_{3p} + i\delta}
\]

where \( E \) is the energy of the photoelectron, \( E_{3p} \) is the energy of a 3p electron. The integral is understood to include a summation over all the discrete nd states.

Since we use a standing-wave representation for the wave function in the continuum, we have \( (\text{ed} \| d \| 3p) = (\text{ed} \| d \| \epsilon d\epsilon) = d_\epsilon \). To simplify the notation, we write

\[
(3pEd \| \Gamma_i(\omega) \| 3p) = \Gamma_i(\omega)
\]

and

\[
(\epsilon d\epsilon Ed \| \Gamma_i(\omega) \| 3p) = G(\omega).
\]

If we express the photon energy \( \omega \) in terms of the energy of the photoelectron \( E \), we can rewrite Eq. (9) as

\[
d_\epsilon = \int d\epsilon \frac{d\epsilon G(\epsilon)}{\omega - \epsilon - i\delta} - \int d\epsilon \frac{d\epsilon G(\epsilon)}{\omega + \epsilon + 2\epsilon}
\]

where \( I_{3p} \) replaces \( E_{3p} \). The small imaginary number \( i\delta \) drops out in the denominator of the last
term of Eq. (10) because this denominator never approaches zero.

\[
\Gamma_{\alpha, E} = (3pE d\| U_1\| \epsilon d 3p) + \int d\epsilon' \frac{(3p\epsilon' d\| U_1\| \epsilon d 3p)\Gamma_{e', E}}{E - \epsilon' - i\delta} = \int d\epsilon' \frac{(3p\epsilon' d\| U_1\| \epsilon d 3p)\Gamma_{e', E}}{E + \epsilon' + 2I_{sp}}
\]

(11)

\[
G_{\alpha, E} = (\epsilon d 3p\| U_1\| 3p\epsilon d)\Gamma_{e', E} = \int d\epsilon' \frac{(\epsilon d 3p\| U_1\| 3p\epsilon d)\Gamma_{e', E}}{E - \epsilon' - i\delta} = \int d\epsilon' \frac{(\epsilon d 3p\| U_1\| 3p\epsilon d)\Gamma_{e', E}}{E + \epsilon' + 2I_{sp}}
\]

(12)

where

\[
(n_i l_i, n_3 l_3\| U_1\| n_2 l_2, n_4 l_4) = \frac{3}{2} (n_i l_i, n_3 l_3\| V_1\| n_2 l_2, n_4 l_4) + \sum_{l_i'} (-1)^{l_i'} \left( l_i \begin{array}{c} 1 \ 1 \ \ l_i' \ l_i' \end{array} \right) (n_i l_i, n_3 l_3\| V_{l_i'}\| n_4 l_4, n_2 l_2).
\]

(13)

The matrix element

\[
(n_i l_i, n_3 l_3\| V_1\| n_2 l_2, n_4 l_4)
\]

is defined in Eq. (16) of Ref. 3.

In the reaction-matrix method, the dipole transition amplitude is written

\[
\left\{ d_E + P \int [d_E \sigma_{e', E} / (E - \epsilon)] d\epsilon \right\} / (1 - i\sigma K_{e', E}),
\]

(14)

where 1/(1 - i\sigma K_{e', E}) is the normalization factor B of Eqs. (1) and (5). The K-matrix elements satisfy

\[
K_{e', E} - V_{e', E} + P \int [V_{e', E} K_{e', E} / (E - \epsilon')] d\epsilon',
\]

(15)

where

\[
V_{e', E} = (3p\epsilon d 1P^0) V_{res} (3p\epsilon d 1P^0)
\]

(16)

with V_{res} given by Eq. (4).

Comparing Eqs. (10) and (14), we see they are very similar, except for the last term of Eq. (10), which does not appear in (14). Thus Eq. (11) differs from Eq. (15) primarily through the occurrence of its last term. This term involves the interaction-matrix element (3p\epsilon d 1P^0) corresponding to virtual excitation of two ground-state particles from 3p to (\epsilon d, \epsilon') d, whereas the other terms depend on interactions that replace one excitation by another one. The inclusion of virtual excitations of pairs of 3p electrons in the RPAE equations means that these equations take ground-state correlations into account while the reaction-matrix method of Ref. 2 does not.

The matrix \( G_{e', E} \) has the main role of representing the effect of these ground-state correlations. Indeed, if we set \( G_{e', E} = 0 \) in Eqs. (10) and (11), then Eq. (10) reduces to Eq. (14) and Eq. (11) reduces to Eq. (15), with \( \Gamma_{e', E} \) and \( K_{e', E} \) related by

\[
\Gamma_{e', E} = \left[ K_{e', E} / (1 - i\sigma K_{e', E}) \right].
\]

(17)

This equivalence is proved by using the identity

\[
\lim(x - i\alpha)^{-1} = P x^{-1} + i\pi\delta(x), \text{ as } \alpha \to 0^+
\]

The integral equations satisfied by \( \Gamma_{e, E} \) and \( G_{e, E} \) are obtained by rewriting Eq. (6) as

In the comparison of the two methods, we have disregarded comparatively minor differences between the matrix elements \( V_{e', E} \) of the reaction-matrix method and the corresponding elements of \( U_{e'} \).

This comparison of RPAE and reaction matrix equations has suggested to us a procedure for incorporating the main effects of ground-state correlations in the reaction-matrix method. This procedure rests on the fact that the matrix element (3p\epsilon d 1P^0)\| U_1\| \epsilon d 3p) for virtual excitation of a pair of ground-state electrons differs from the matrix element (3p\epsilon d 1P^0)\| U_1\| \epsilon d 3p) for replacement of a single excitation by an amount that is small in the situations where pair excitations are important. The difference lies entirely in the \( \sum_{l_i'} \) terms of Eq. (13), whereas the first term is dominant and identical for the two processes. If we set

\[
(3p\epsilon d 1P^0)\| U_1\| \epsilon d 3p) \approx (3p\epsilon d 1P^0)\| U_1\| \epsilon d 3p)
\]

(18)

in Eqs. (11) and (12), these equations become symmetric in \( \Gamma_{e, E} \) and \( G_{e, E} \) and are solved by setting

\[
\Gamma_{e, E} = G_{e, E},
\]

(19)

where \( \Gamma_{e, E} \) satisfies

\[
\Gamma_{e, E} = (3pE d\| U_1\| \epsilon d 3p)
\]

\[
+ \int d\epsilon' \left[ (E - \epsilon' - i\delta)^{-1} - (E + \epsilon' + 2I_{sp})^{-1} \right]
\]

\[
\times (3p\epsilon' d\| U_1\| \epsilon d 3p)\Gamma_{e', E}.
\]

(20)

To estimate the inaccuracy of Eq. (18), we note that the first term of Eq. (13) is proportional to the \( 9^1 \) integral of Fano and Cooper which is a generalization of Slater's \( G^1 \) integral to off-diagonal matrix elements. The coefficient in front of the \( 9^1 \) integral is the largest among all the terms in Eq. (13). The term with \( 9^1 \) represents the exchange interaction between the excited electron and all the electrons that remain in the ground-state shell 3p. As the numerical calculation of Starace shows, the \( 9^1 \) integral contributes more
than 85% of the magnitude of all the residual interaction-matrix elements of Ar and more than 80% for those of Xe.

On the basis of the similarity between Eqs. (20) and (15), we propose that for a single channel problem, the ground-state correlation effect be included in an *ad hoc* way by solving

\[
\tilde{R}_{\epsilon, \epsilon'} = V_{\epsilon, \epsilon'} + P \int [V_{\epsilon', \epsilon''} \tilde{R}_{\epsilon, \epsilon''} / (E - \epsilon')] d\epsilon'
\]

\[
- \int [V_{\epsilon, \epsilon'} \tilde{R}_{\epsilon', \epsilon''} / (E + \epsilon + 2I_{sp})] d\epsilon'
\]

(21)

instead of Eq. (15). The dipole transition amplitude is then obtained by calculating

\[
\left( d_{\epsilon} + P \int [d_{\epsilon} \tilde{R}_{\epsilon, \epsilon''} / (E - \epsilon)] d\epsilon 
\right)
\]

\[
- \int [d_{\epsilon'} \tilde{R}_{\epsilon, \epsilon''} / (E + \epsilon + 2I_{sp})] d\epsilon
\]

\[
\times (1 - i\pi \tilde{R}_{\epsilon, \epsilon''})^{-1}
\]

(22)

instead of Eq. (14). We call our proposed method SRPAE because Eqs. (21) and (22) are obtained by approximating the RPAE equations. The calculation is to be done entirely by the procedure of the reaction-matrix method. The advantage of this approach is that one can obtain a much better numerical result, as will be shown in Sec. III, without spending as much computational time as in the RPAE method. The importance of the ground-state correlation effect can be evaluated by comparing the results calculated by SRPAE and by the reaction-matrix method.

The main Eqs. (21) and (22) of SRPAE rest on the approximate Eq. (18), whose range of validity is discussed in the paragraph following Eq. (20). Equations (21) and (22) are similar to Wendin’s Eqs. (3.7)–(3.9) of Ref. 4, except for a minor difference in the definition \( V_{\epsilon, \epsilon'} \). By analyzing the polarization propagator diagrammatically, Wendin argues that the interaction-matrix elements for the final state and the ground state are exactly the same,\(^11\) that is, Eq. (18) holds rigorously. We have assumed Eq. (18) to hold only when the \( g^1 \) integral dominates. In any actual calculations, the magnitude of the residual interaction-matrix elements for the final state and for the ground state depends on the choice of basis set. Whether or not Eq. (18) holds probably depend on the basis set also.

### III. RESULTS AND CONCLUSIONS

The numerical procedure to calculate the photoionization cross section by the SRPAE method is basically the same as that used in the reaction-matrix method, except for replacing Eqs. (14) and (15), respectively, by Eqs. (22) and (21). One starts by generating the single-particle wave function from the Schrödinger equation using Herman-Skillman potential. These wave functions are used to evaluate the dipole-matrix elements \( d_{\epsilon} \) and the residual interaction-matrix elements \( V_{\epsilon, \epsilon'} \). Then one solves the integral Eq. (21). The dipole amplitude is obtained by doing the integrations in Eq. (22).

The method for evaluating the residual interaction-matrix elements \( V_{\epsilon, \epsilon'} \) and for solving the integral Eq. (21), which involves the principal-part integration, have been discussed by Altick and Moore\(^12\) and by Starace.\(^2\) In this paper, we intend to use the matrix elements \( V_{\epsilon, \epsilon'} \) and dipole amplitudes \( d_{\epsilon} \) calculated by Starace\(^2\) for the \( 3p \rightarrow Ed \) transition of Ar and \( 4d \rightarrow Ef \) transition of Xe. In fact, it was found feasible and most expedient to introduce a further numerical approximation\(^13\) for \( V_{\epsilon, \epsilon'} \), namely, by factorizing it in the form

\[
V_{\epsilon, \epsilon'} = g(\epsilon)g(\epsilon')
\]

(23)

No general arguments supporting the validity of this assumption seem to have been advanced, but the numerical values of \( V_{\epsilon, \epsilon'} \) calculated by Starace, are approximated by Eq. (23) to within 5%. It should be noted that the approximation (23) appears to be valid under the same circumstances as (18), that is, when the \( g^1 \) integral of Ref. 10 is dominant. In fact the quasifactorability of \( g^1 \) and its physical origin were discussed on p. 486 of this reference.

The main effect of the factorization (23), also noted in Ref. 13, is that it permits an analytical solution of the reaction-matrix equation. Substituting (23) into Eq. (21) we find

\[
\tilde{R}_{\epsilon, \epsilon'} = V_{\epsilon, \epsilon'} / \left(1 - P \int V_{\epsilon, \epsilon''} / (E - \epsilon) d\epsilon + \int V_{\epsilon', \epsilon''} / (E + \epsilon + 2I_{sp}) d\epsilon \right)
\]

(24)

for the \( \tilde{R} \) matrix elements. The corresponding \( K \)-matrix element in the factorizable approximation has the same form, except for the absence of the last term in the denominator of Eq. (24).

The photoionization cross section calculated by the SRPAE method using Eqs. (24) and (22) are shown on Fig. 1 for the \( 3p \rightarrow Ed \) transition of Ar, and on Fig. 2 for the \( 4d \rightarrow Ef \) transition of Xe. The results of the calculations by Starace\(^2\) using the reaction-matrix method are also given. By comparing with the experimental results of Samson\(^14\) for Ar and of Haensel et al.\(^15\) for Xe, we find that the cross sections calculated by the SRPAE method are much better than those obtained by Starace without any ground-state correlation. The improvement is substantial, especial-
ly on the higher-energy side. Note that the results shown on the figures for both calculations include the transition $3p - Ed$ of Ar and $4d - Ef$ of Xe only, whereas the experiments measure the total cross sections. The agreement with experiment of the calculation by SRPAE method in Ar is not as good as for Xe. This is probably due to the limitations of RPAE itself, which assumes in essence a high density of electrons in the shell to be ionized; this density is higher for the $4d^{10}$ shell of Xe than for the $3p^6$ shell of Ar.

A comparison of the SRPAE results with the full RPAE calculation by Amusia et al. is possible. We plot in Fig. 1 the result of the RPAE calculation for Ar, but not on Fig. 2 for Xe, because in the latter case the RPAE results of Amusia et al. agree with our SRPAE results to within 5–10% over the energy range considered. For Ar, the results of the two calculations deviate from experiment by approximately the same amount, but lie on opposite sides of the experimental curve. This probably results from differences in the interaction-matrix elements used in the two calculations.

We conclude that the SRPAE method provides a simple and effective way to calculate the photoionization cross section with simultaneous improvement of the ground-state and final-state wave functions. Even though this method probably cannot give better results than the complete RPAE calculation, the simplicity of its numerical procedure should be emphasized; the SRPAE appears to provide a simple way to calculate the dipole transition amplitudes rather accurately when the independent-particle approximation fails. The SRPAE should be applicable as long as the $g^1$ integral dominates the matrix elements of the residual interaction. Generalization of this method to two or more channels is possible, but one has to check the validity of the approximation first.

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1 U. Fano and J. W. Cooper, Rev. Mod. Phys. 40, 441 (1968), Sec. 4 and references therein.
9 RPA is ordinarily used in high-density media and employs a Hartree zeroth approximation. RPAE, as used by Amusia et al., presupposes a Hartree–Fock zeroth approximation and is a generalization of RPA.
10 Reference 1, Sec. 6.1.
11 In Ref. 4, Wendin also considers a modification of his Eqs. (3.5) and (3.9) that reduces the ground-state interaction matrix element by a factor of \((N-2)/(N-1)\), where \(N\) is the number of electrons in a closed shell.
13 Reference 1, Sec. 6.3.